

UVEX: A Time Domain Discovery Machine

UV Photometry as a Probe of Stellar Explosions

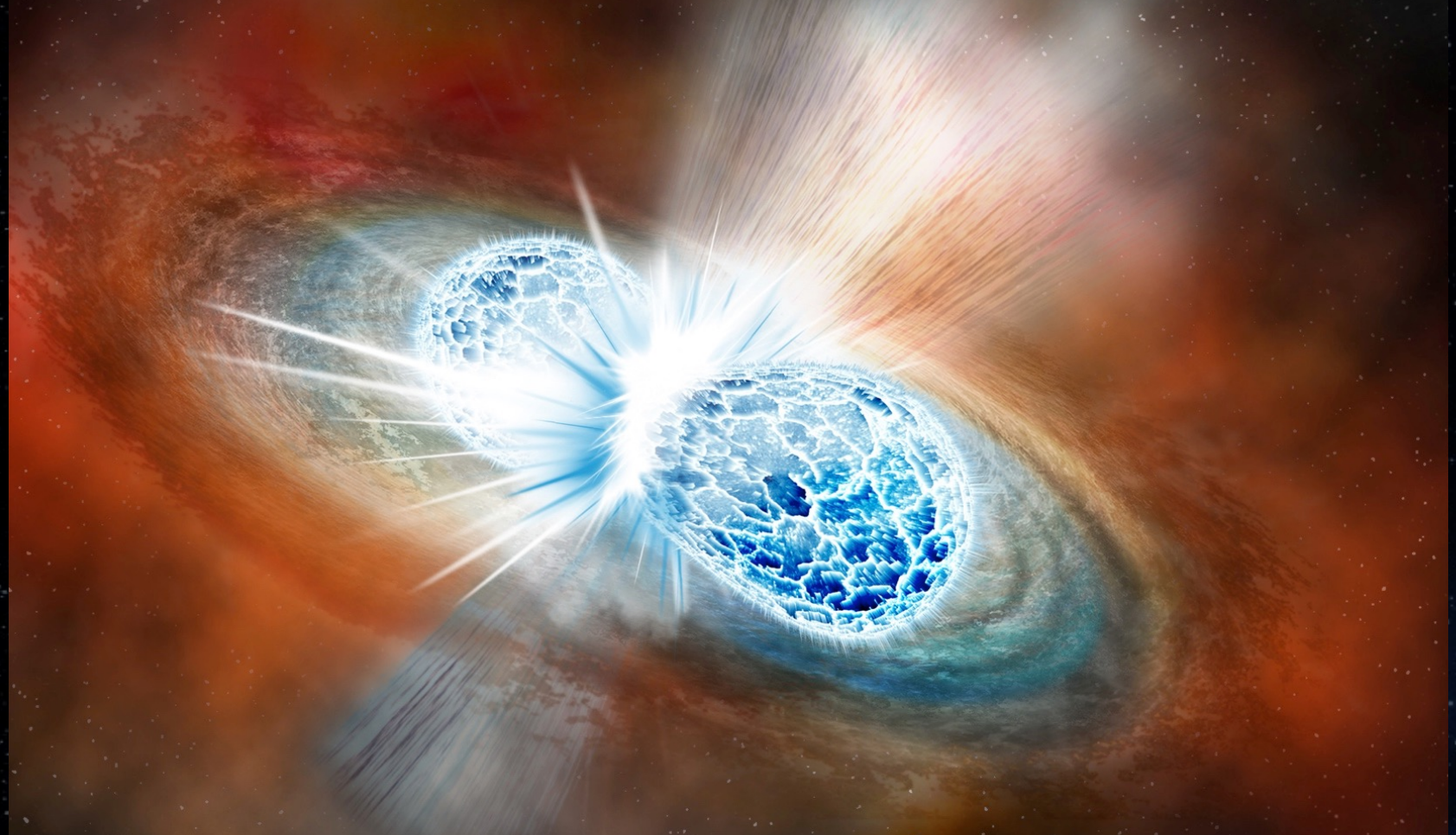
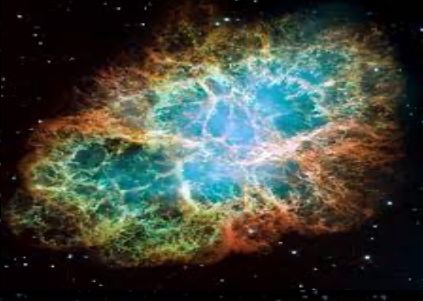


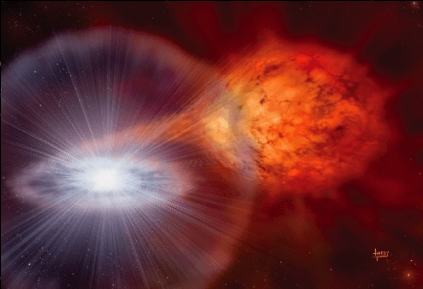
Image Credit: Robin Dienel/Carnegie Observatories

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UV Photometry as a Probe of Stellar Explosions



Core-Collapse Supernova



Thermonuclear Supernova



Peculiar Transients

Why (broadly) the ultraviolet?

Stellar explosions often start hot and then rapidly cool

Why (broadly) the ultraviolet?

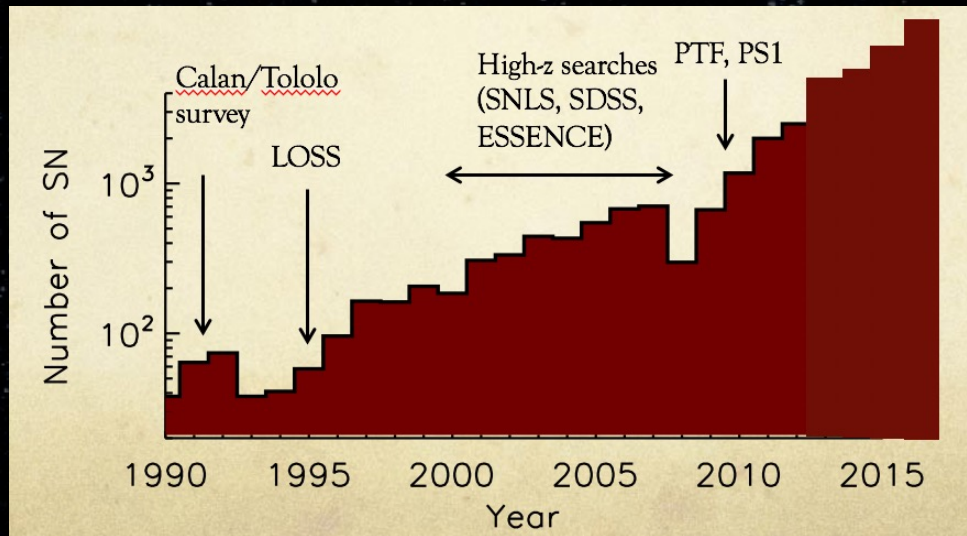
Stellar explosions often start hot and then rapidly cool

- UV is prime a prime wavelength to discover young transients

Why (broadly) the ultraviolet?

Stellar explosions often start hot and then rapidly cool

- UV is prime a prime wavelength to discover **young** transients

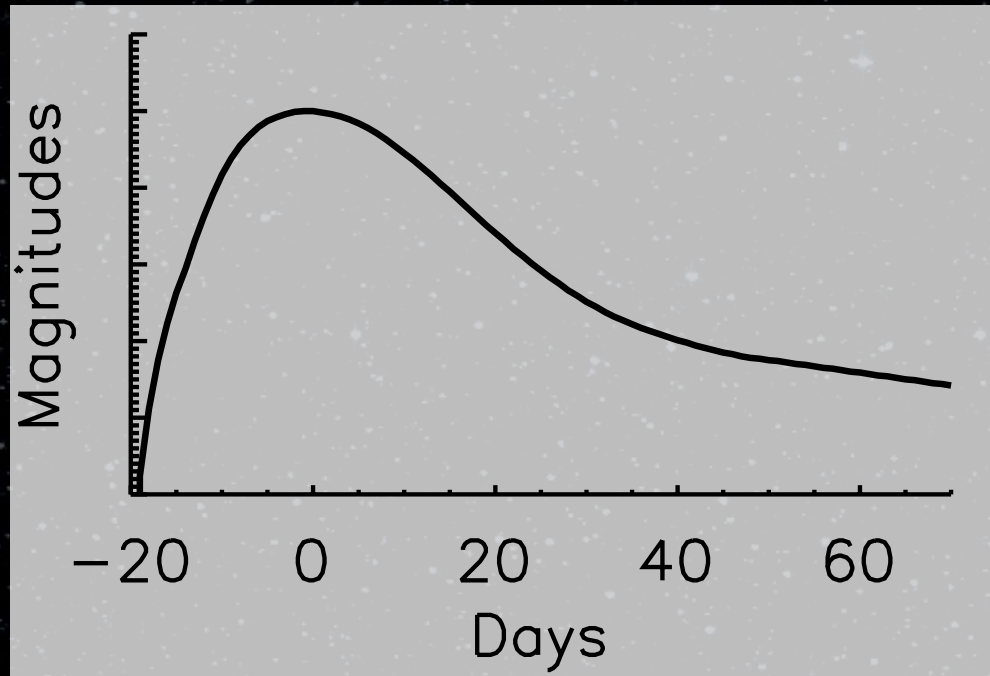


In current era of optical time domain supernova, the community discovers >20,000 explosive transients per year

Why (broadly) the ultraviolet?

Stellar explosions often start hot and then rapidly cool

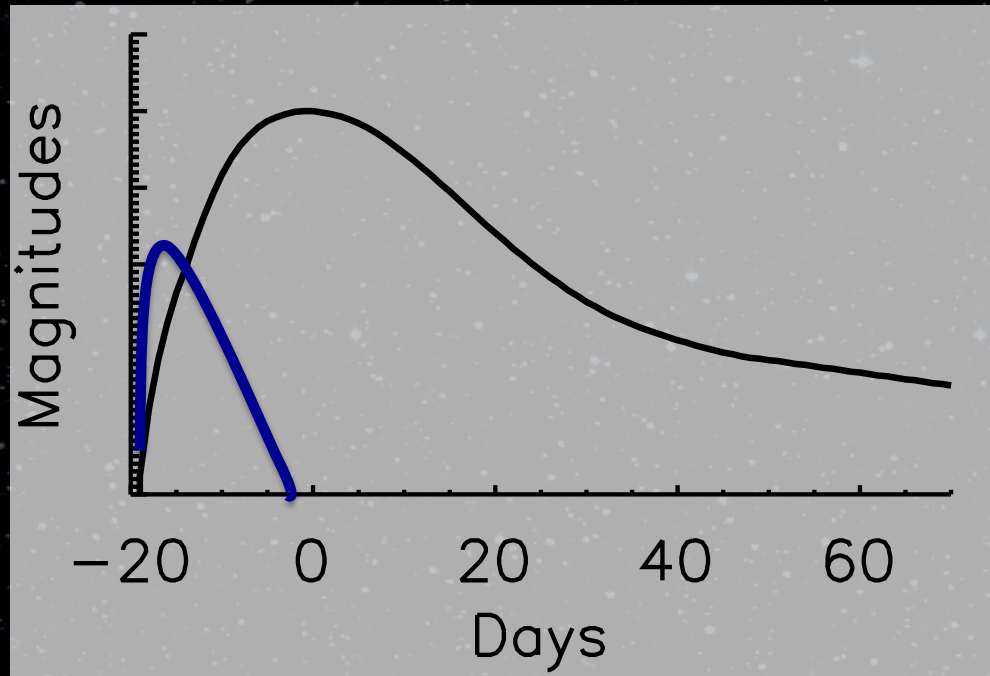
- UV is prime a prime wavelength to discover young transients
- Early UV emission can probe unique physics.



Why (broadly) the ultraviolet?

Stellar explosions often start hot and then rapidly cool

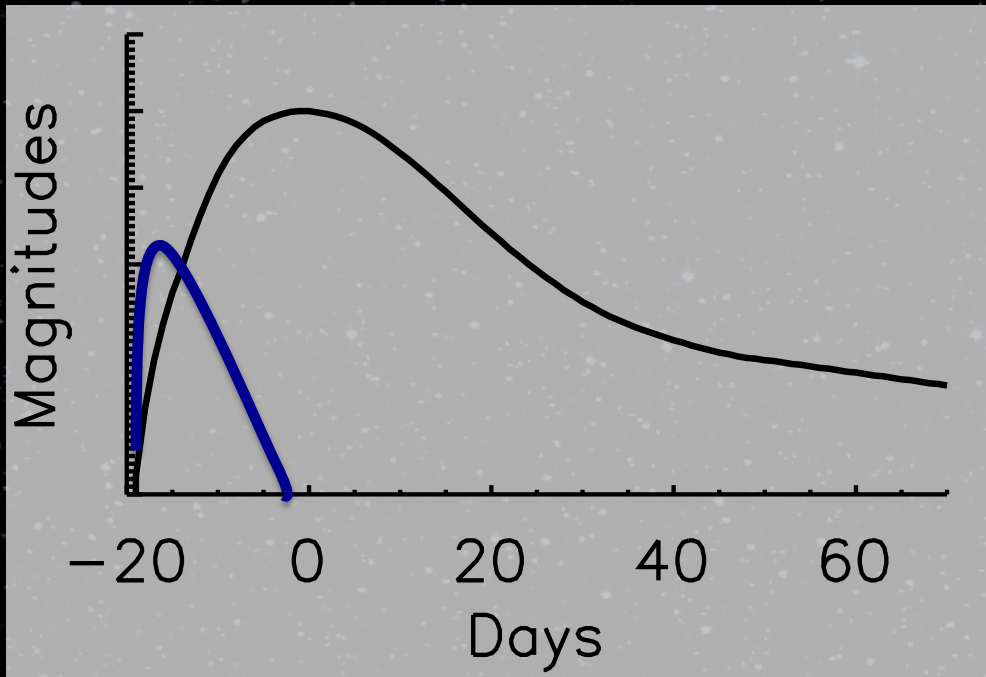
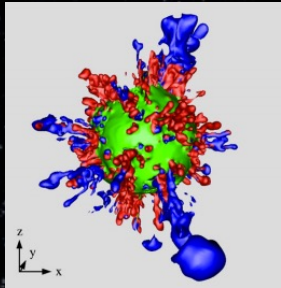
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Why (broadly) the ultraviolet?

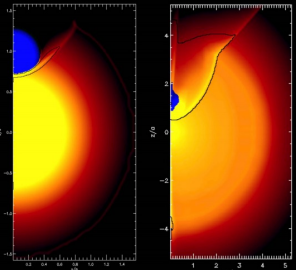
Stellar explosions often start hot and then rapidly cool

- UV is prime a prime wavelength to discover young transients
- Early UV emission can probe unique physics.



Early Emission from:

- Envelope cooling
- Collision with a companion star
- Mixing of radioactive material
- ... etc.



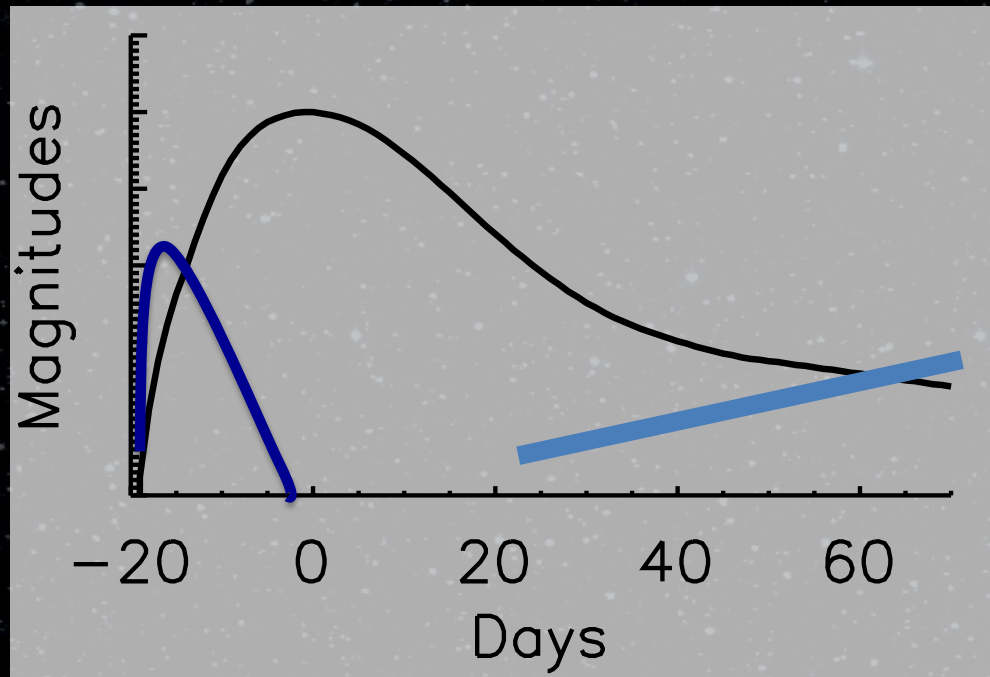
Kasen (2010)

Hammer+ (2010)

Why (broadly) the ultraviolet?

Stellar explosions often start hot and then rapidly cool

- UV is prime a prime wavelength to discover young transients
- Early (and late!) UV emission can probe unique physics.

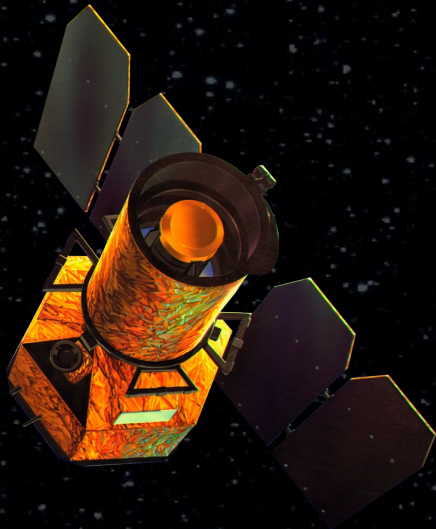


Late-time emission:

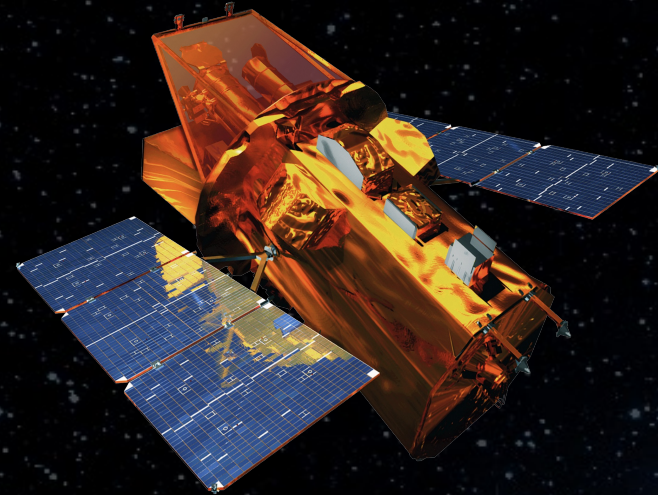
- Interaction with ambient material.
- Additional power from a compact object (accretion, PWN)

What is the current landscape?

Probing Early UV Emission of explosive transients requires wide-field monitoring and/or rapid-response capabilities



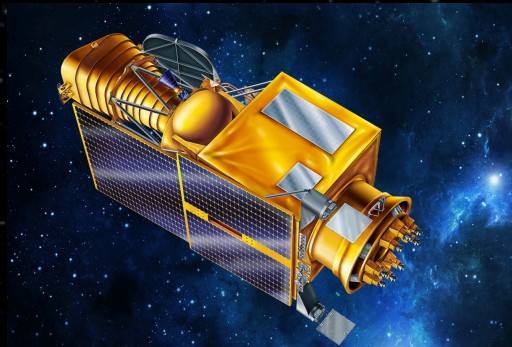
GALEX Time Domain Survey
40 deg², ~2 day cadence,
24 total epochs
(Gezari et al. 2013,2015a,b)



Swift-UVOT ToOs
Hundreds of SN light curves
of date in response to
external triggers

What is the current landscape?

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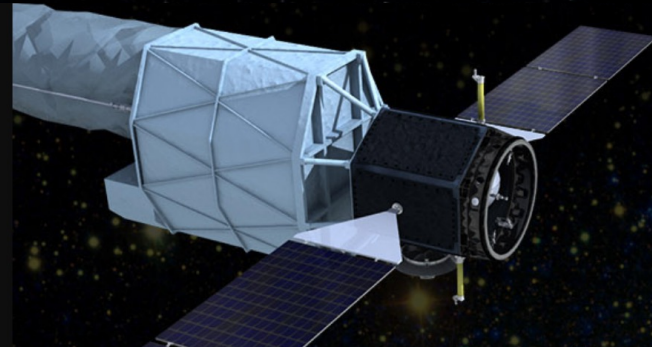


UltraSAT

200 deg²

8.3" FWHM

22.4 mag (900s)

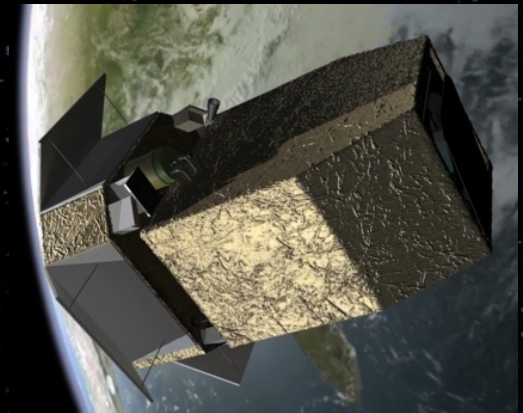


UVEX

10 deg²

2.5" FWHM

24.5 mag (900s)



CASTOR

0.25 deg²

0.15" FWHM

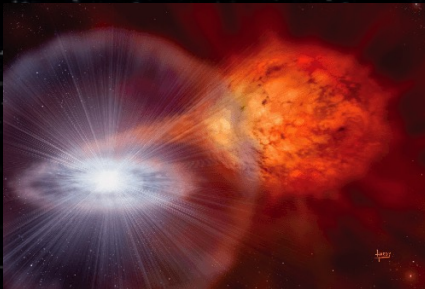
27.2 mag (900s)

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Peculiar Transients

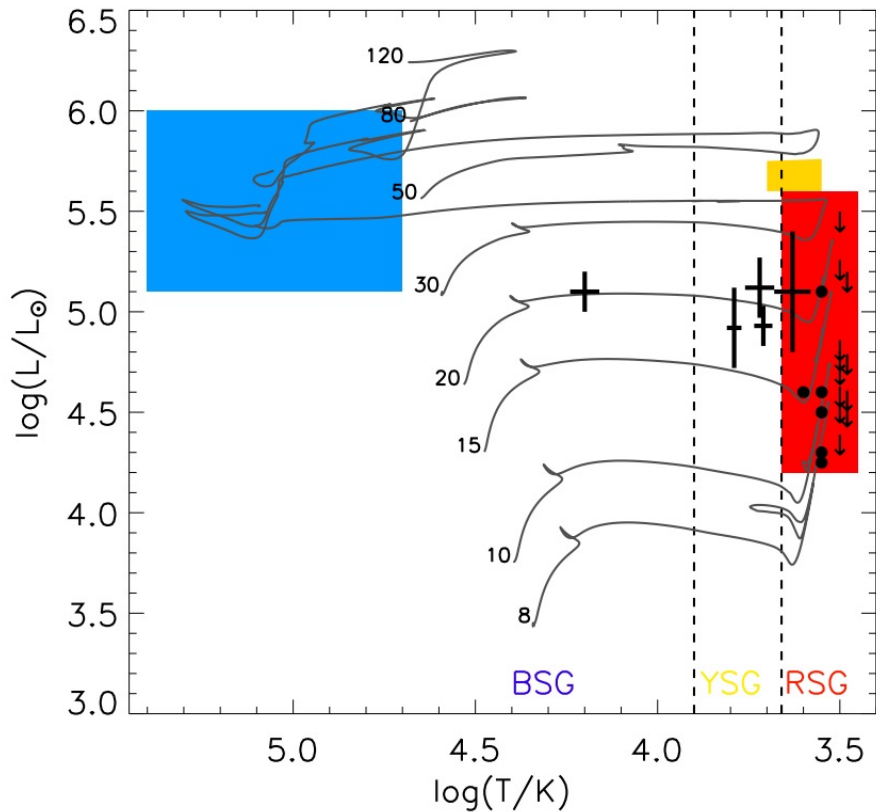
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Core Collapse Supernovae: Science Themes

What are the progenitors to core-collapse supernova?
What is their final structure and final behavior while
evolving towards core-collapse?

UVEX: A Time Domain Discovery Machine

Core Collapse Supernovae: Science Themes



Supernova studies have found progenitors we did not expect, a lack of progenitors that were anticipated, and pointed to classes of stars that we have yet to observationally identify.

Eldridge, et al. (2013)

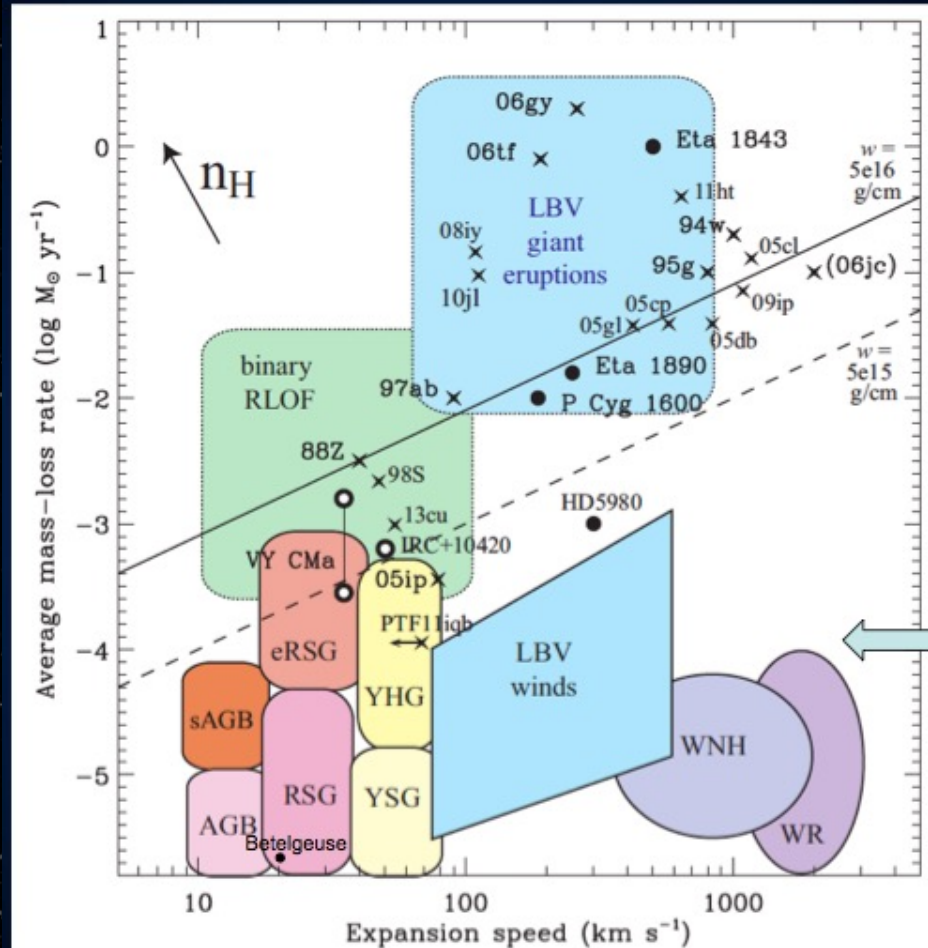
UVEX: A Time Domain Discovery Machine

Core Collapse Supernovae: Science Themes

At this point, it is clear that many massive stars lead dynamic lives during their final years.

UVEX: A Time Domain Discovery Machine

Core Collapse Supernovae: Science Themes

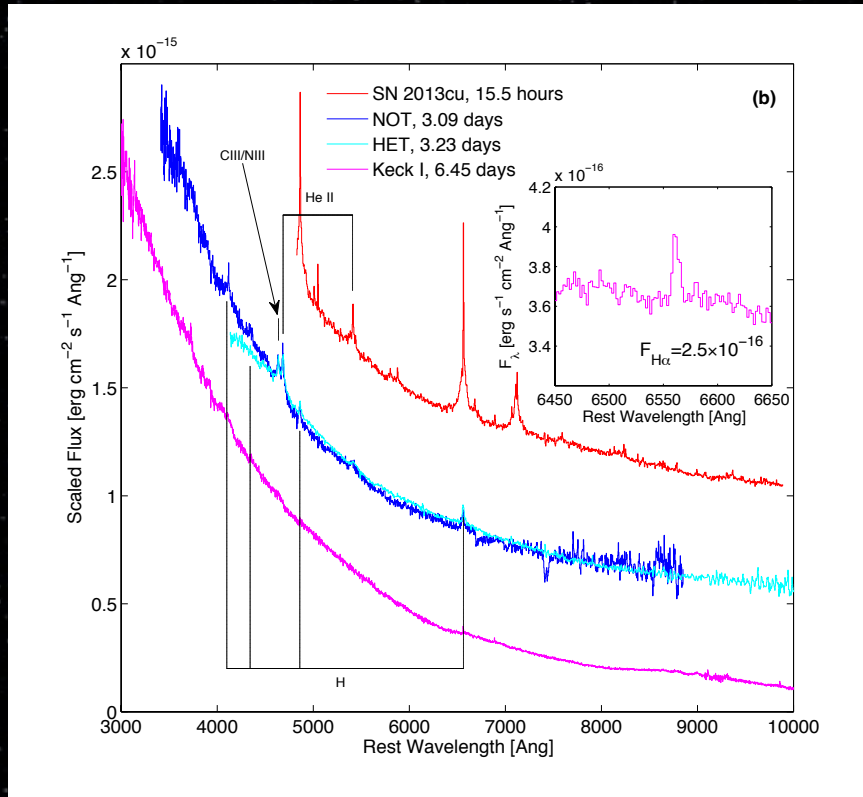


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Exhibit A: Type IIa Supernovae

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Core Collapse Supernovae: Science Themes

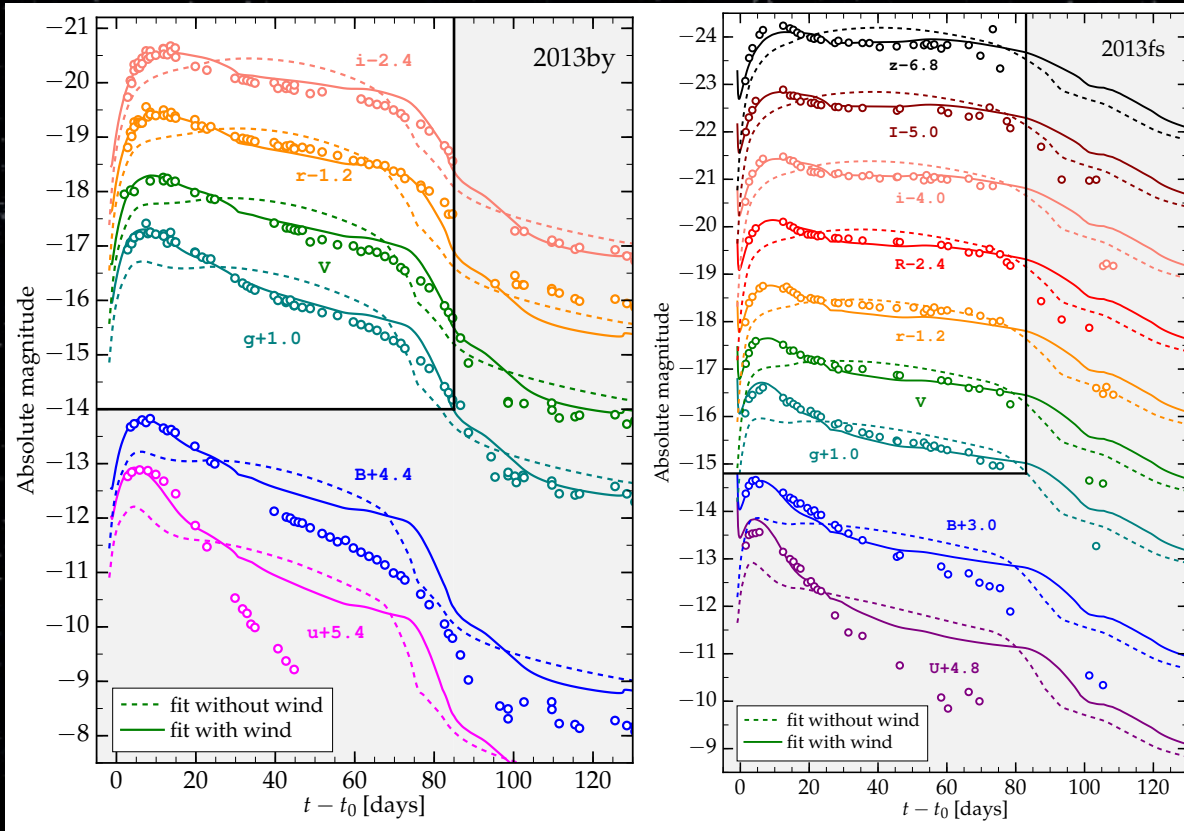


At this point, it is clear that many massive stars lead dynamic lives during their final years.

Exhibit B: enhanced mass loss in “normal supernovae”

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Core Collapse Supernovae: Science Themes

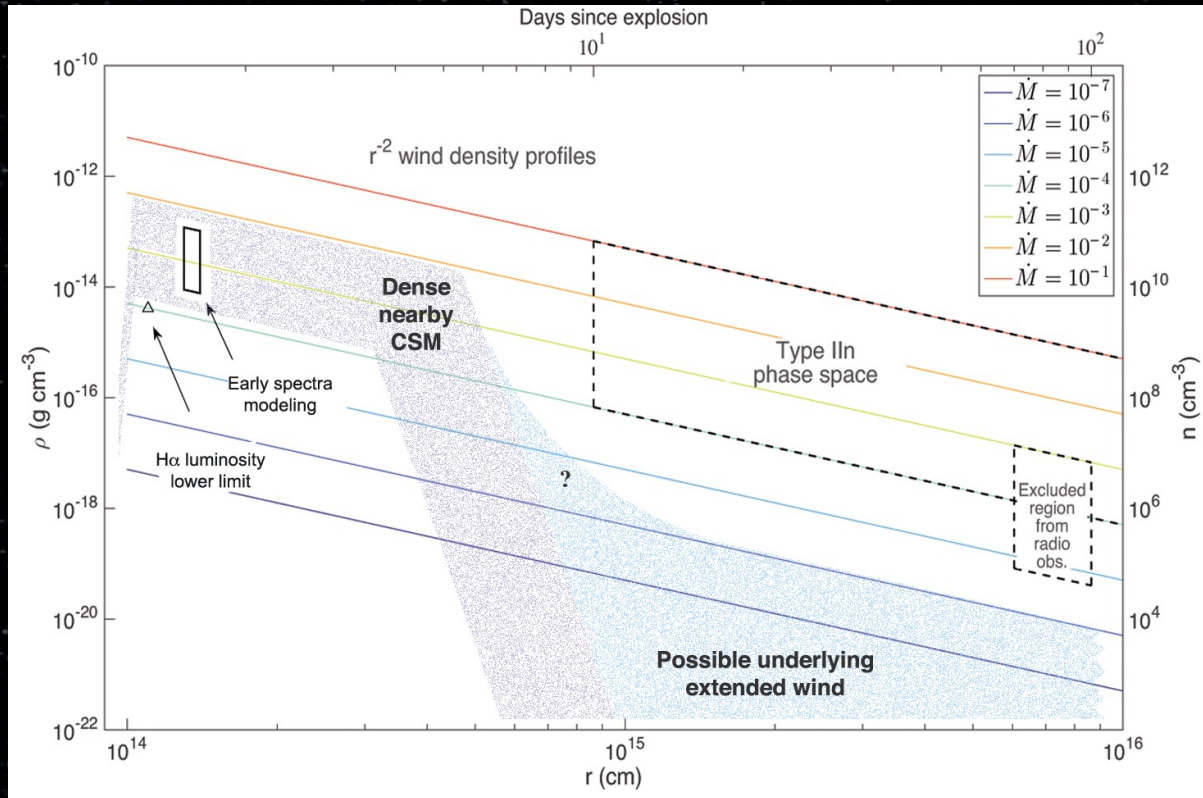


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Exhibit B: enhanced mass loss in “normal supernovae”

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Core Collapse Supernovae: Science Themes



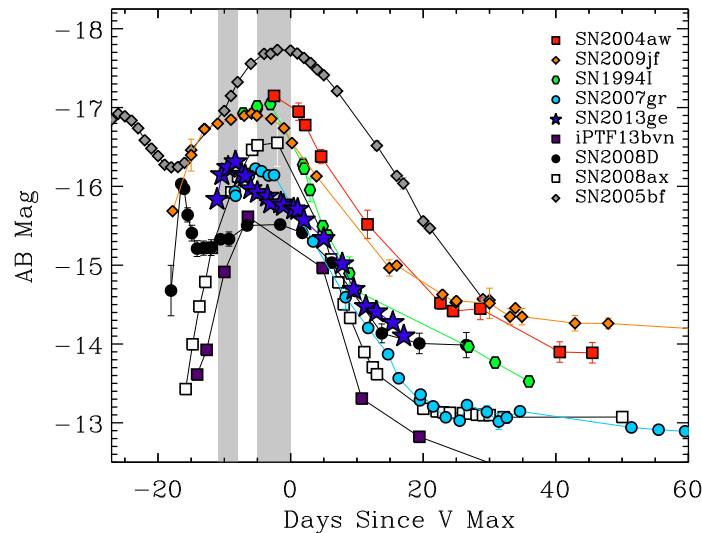
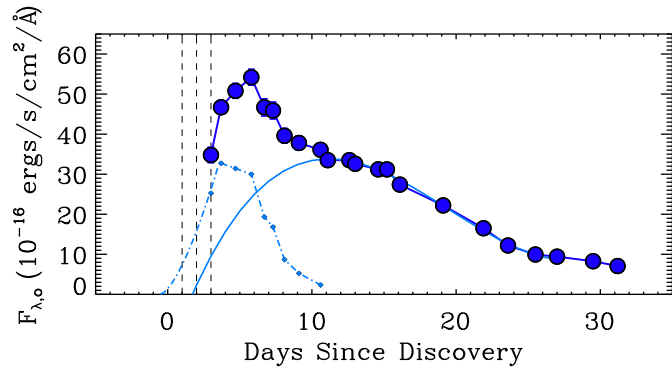
At this point, it is clear that many massive stars lead dynamic lives during their final years.

Exhibit B: enhanced mass loss in “normal supernovae”

Yaron et al. 2017.

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Core Collapse Supernovae: Science Themes

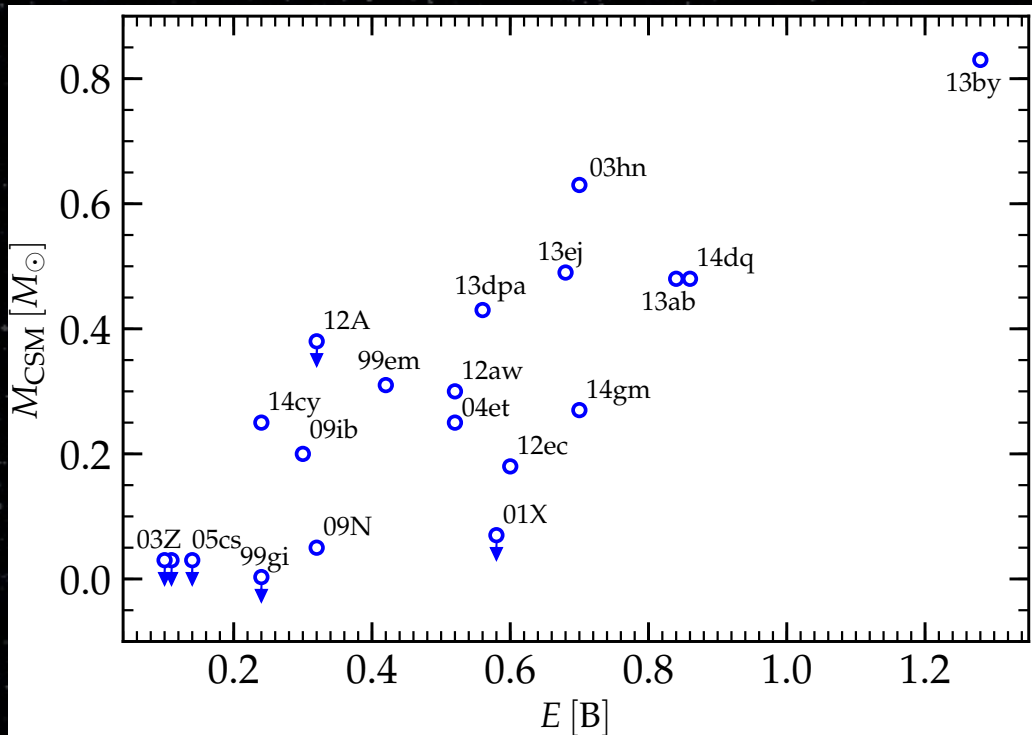


At this point, it is clear that many massive stars lead dynamic lives during their final years.

Exhibit C: This is not limited to hydrogen-rich supernovae

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Core Collapse Supernovae: Science Themes



Morozova, Piro, & Valenti 2018
see also Kazov et al. 2016, Hozzeinzadeh et al. 2018

At this point, it is clear that many massive stars lead dynamic lives during their final years.

How common is this?

Current estimates range from 30-70%

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Core Collapse Supernovae: Science Themes



Fuller et al.

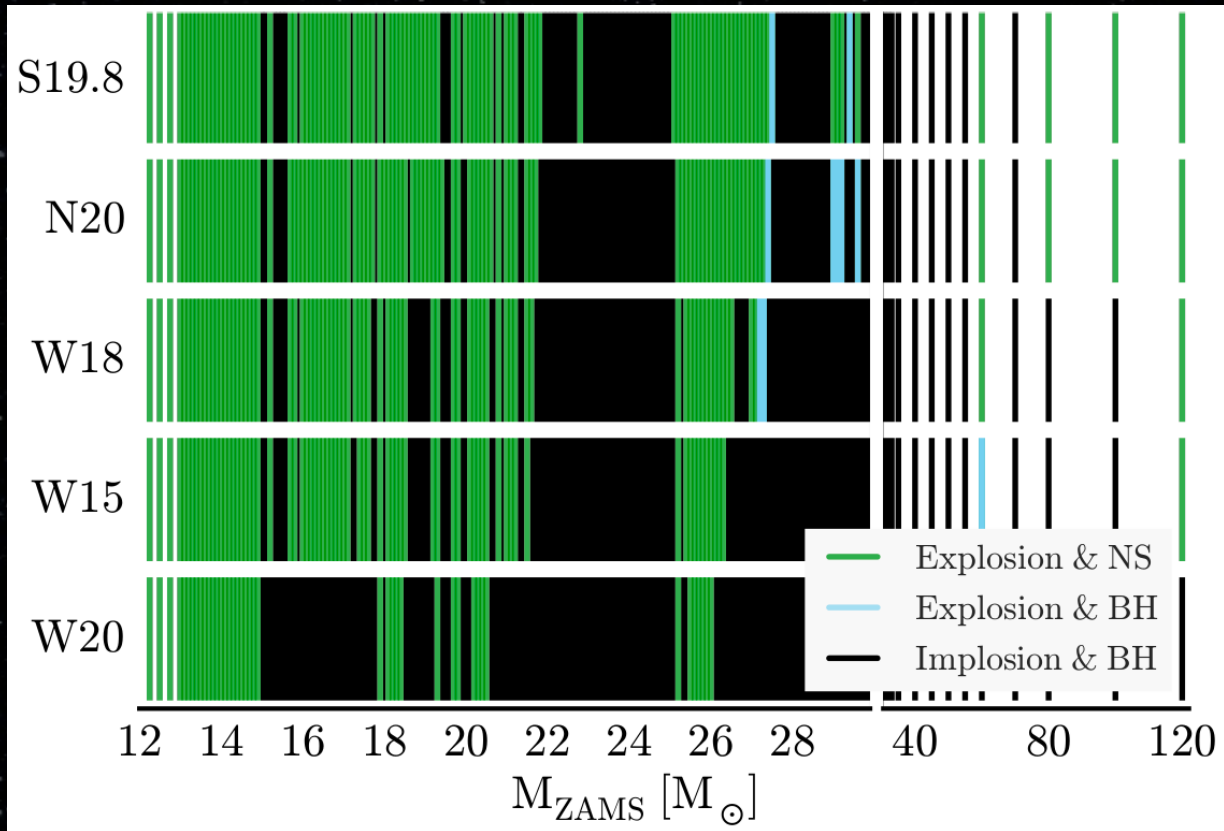
At this point, it is clear that many massive stars lead dynamic lives during their final years.

Why does this matter?

Can impact pre-explosion structure of the star and spin rate of the core

UVEX: A Time Domain Discovery Machine

Core Collapse Supernovae: Science Themes



At this point, it is clear that many massive stars lead dynamic lives during their final years.

Why does this matter?

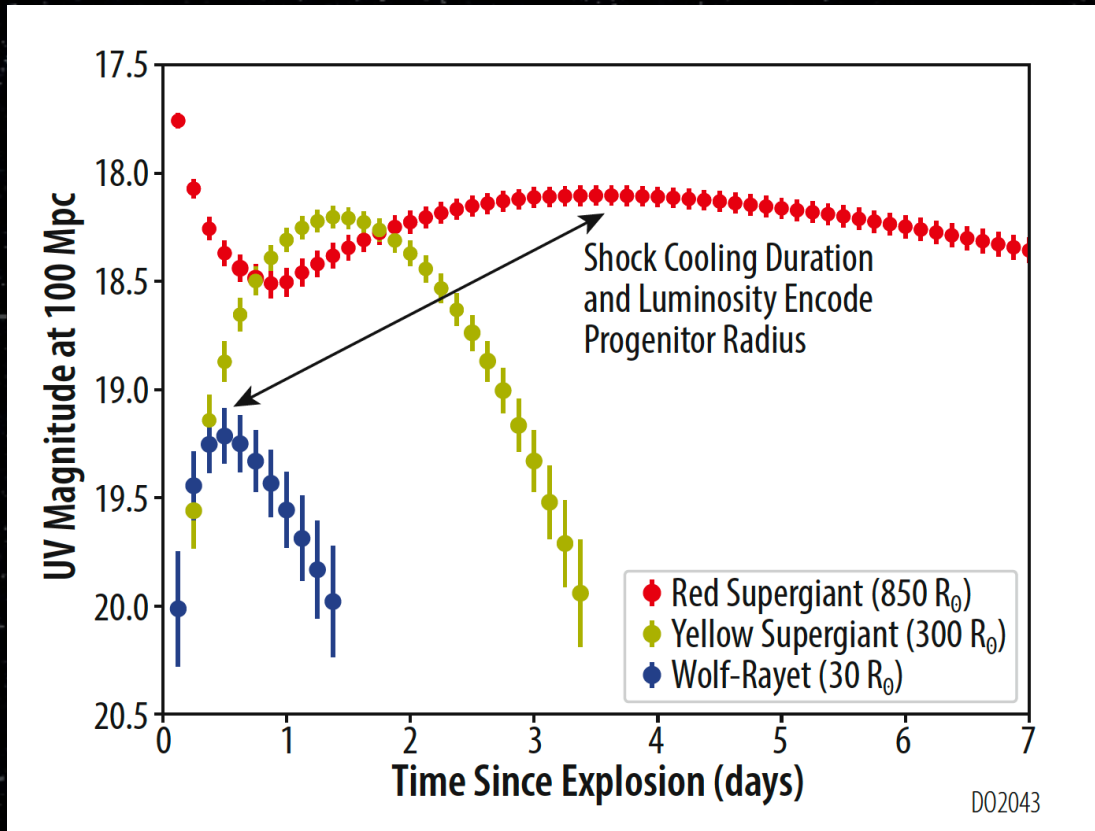
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Core Collapse Supernovae: The Role of UV Data

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Core Collapse Supernovae: The Role of UV Data

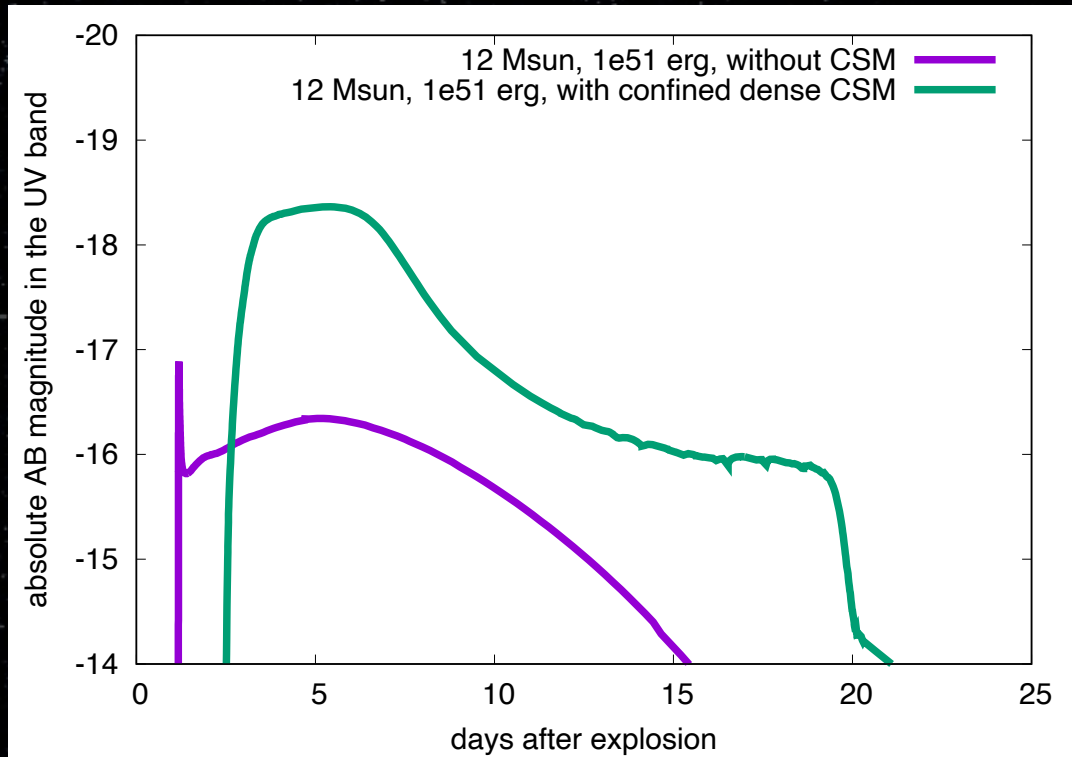


Progenitor Radius
Constraints from
shock cooling
emission.

Figure: Piro, Cenko, Dorado team

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Core Collapse Supernovae: The Role of UV Data



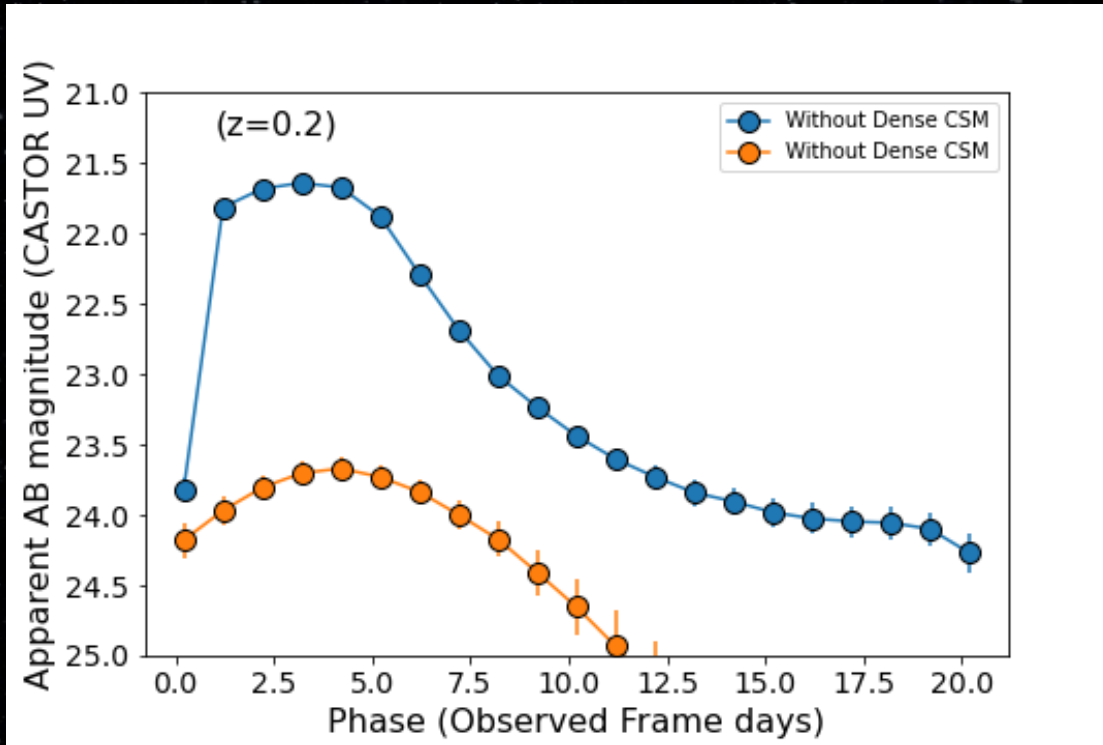
Progenitor Radius
Constraints from
shock cooling
emission.

Fraction of SN with
enhanced pre-SN
mass loss

Figure: T. Moriya

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Core Collapse Supernovae: The Role of UV Data



Progenitor Radius
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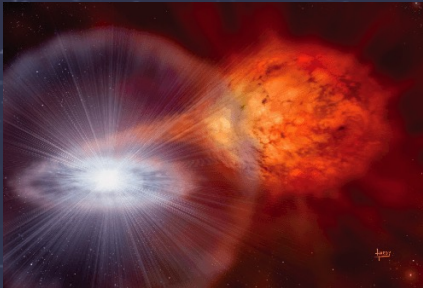
Figure: Drout, CASTOR team

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Thermonuclear Supernova

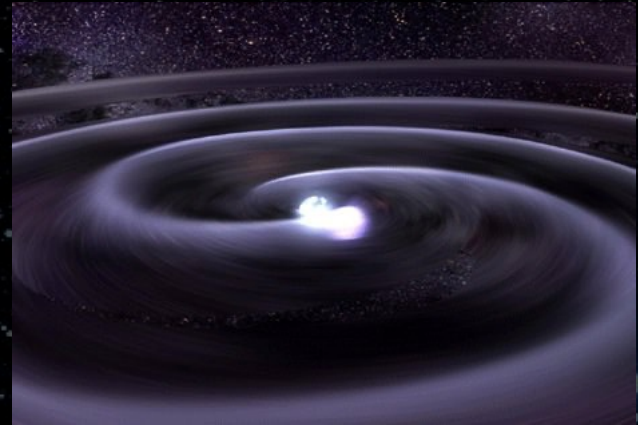
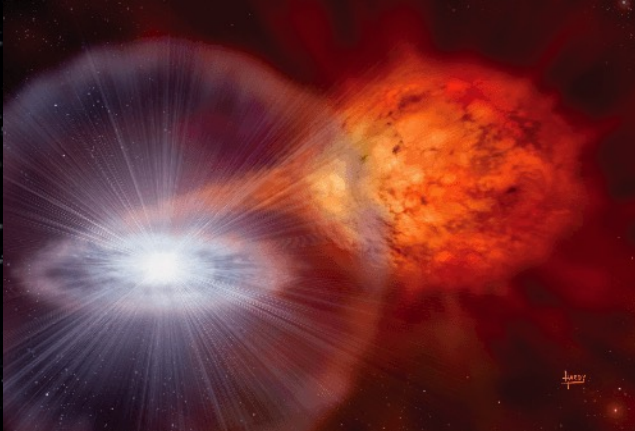


Peculiar Transients

UVEX: A Time Domain Discovery Machine

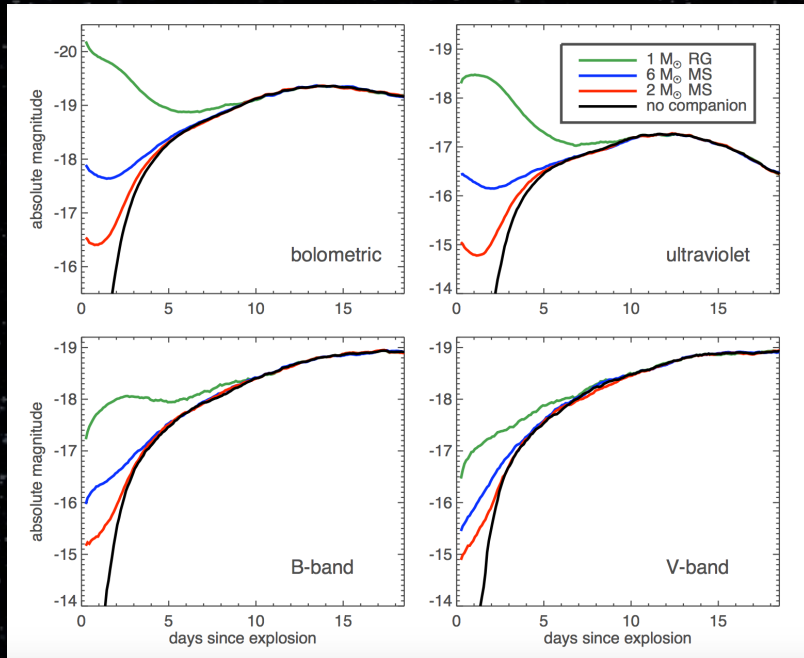
Thermonuclear Supernovae: Science Themes

What is the nature of the binary system? What is the explosion mechanism

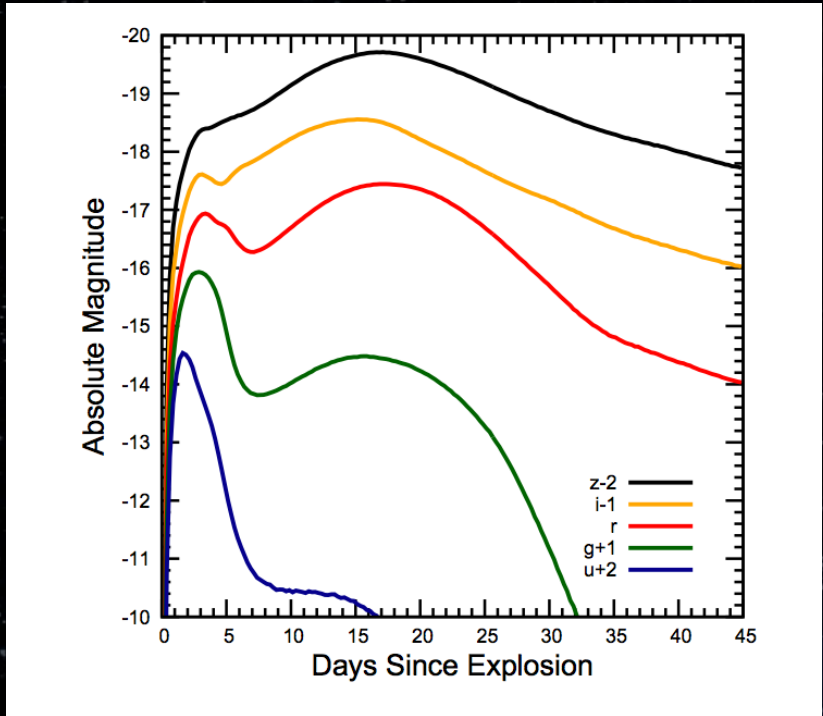


UVEX: A Time Domain Discovery Machine

Thermonuclear Supernovae: Science Themes



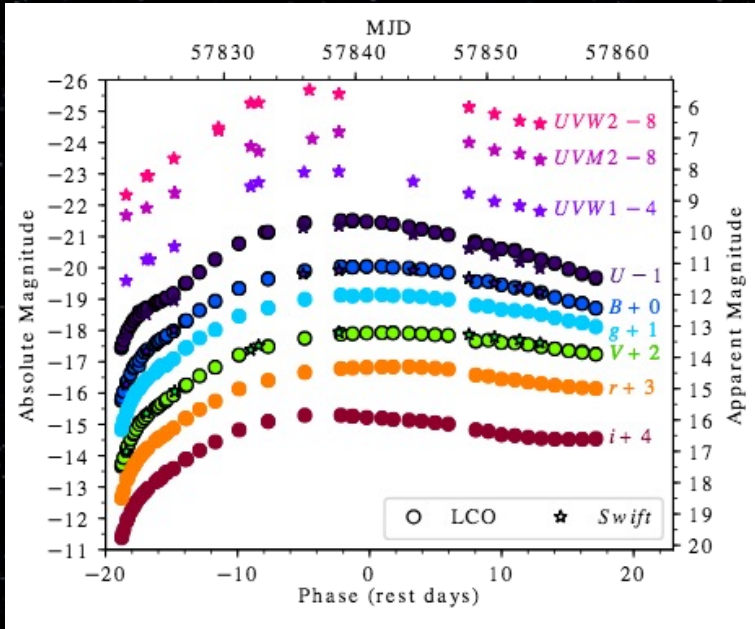
Kasen (2010) Collision with a companion



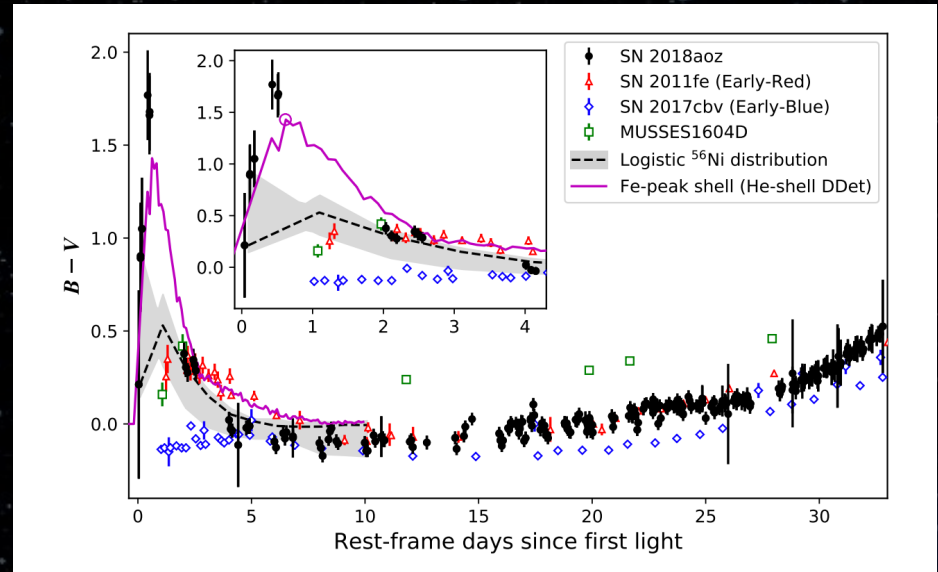
Polin+ (2019) Surface Radioactive decay due to a helium shell detonation

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Thermonuclear Supernovae: Science Themes



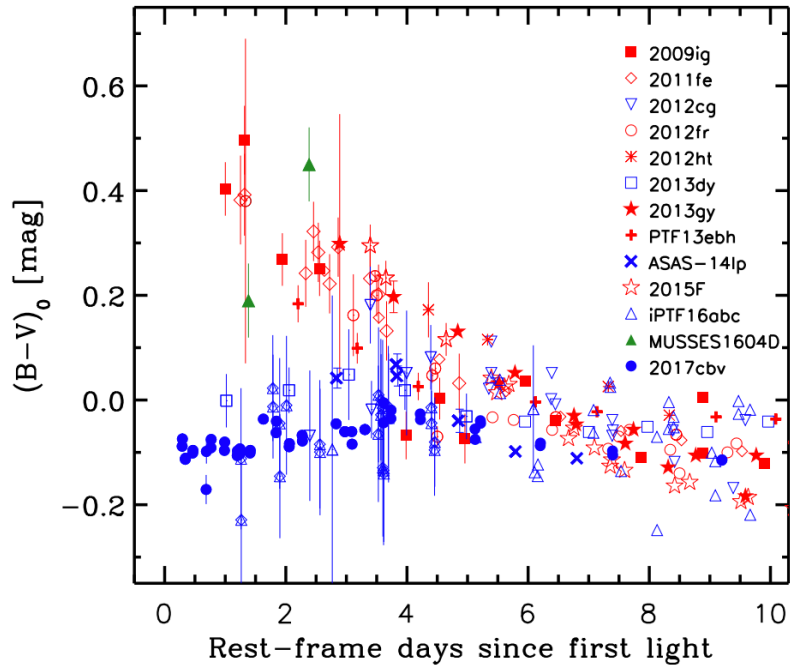
Hosseinzadeh et al (2017)
Type Ia with a blue bump



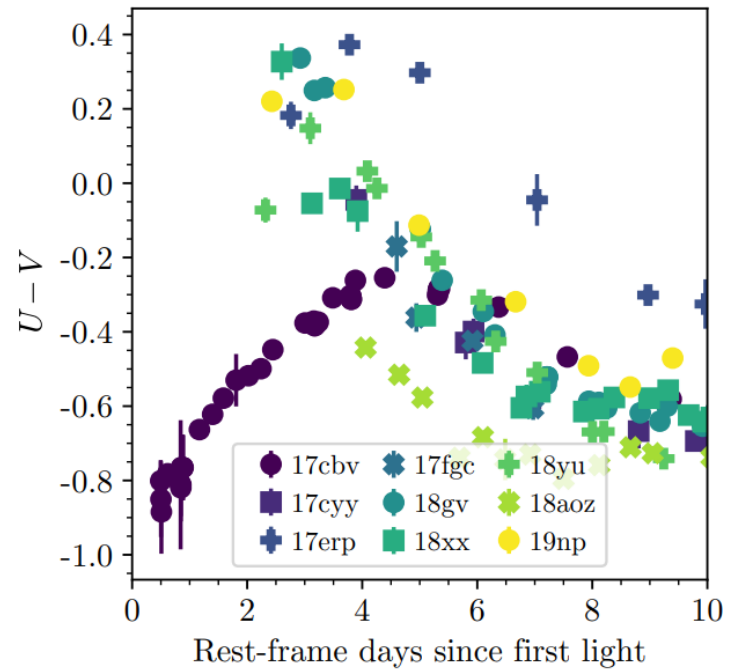
Ni, Moon, Drout et al+ (2022) Type Ia with early excess emission and reddening due to surface Nickel.

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Thermonuclear Supernovae: Science Themes



Stritzinger et al. 2018



Burke et al. (2022)



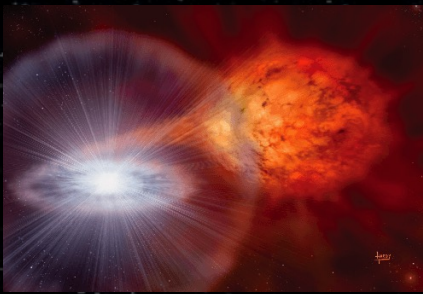
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Thermonuclear Supernovae: The Role of the UV

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Peculiar Transients

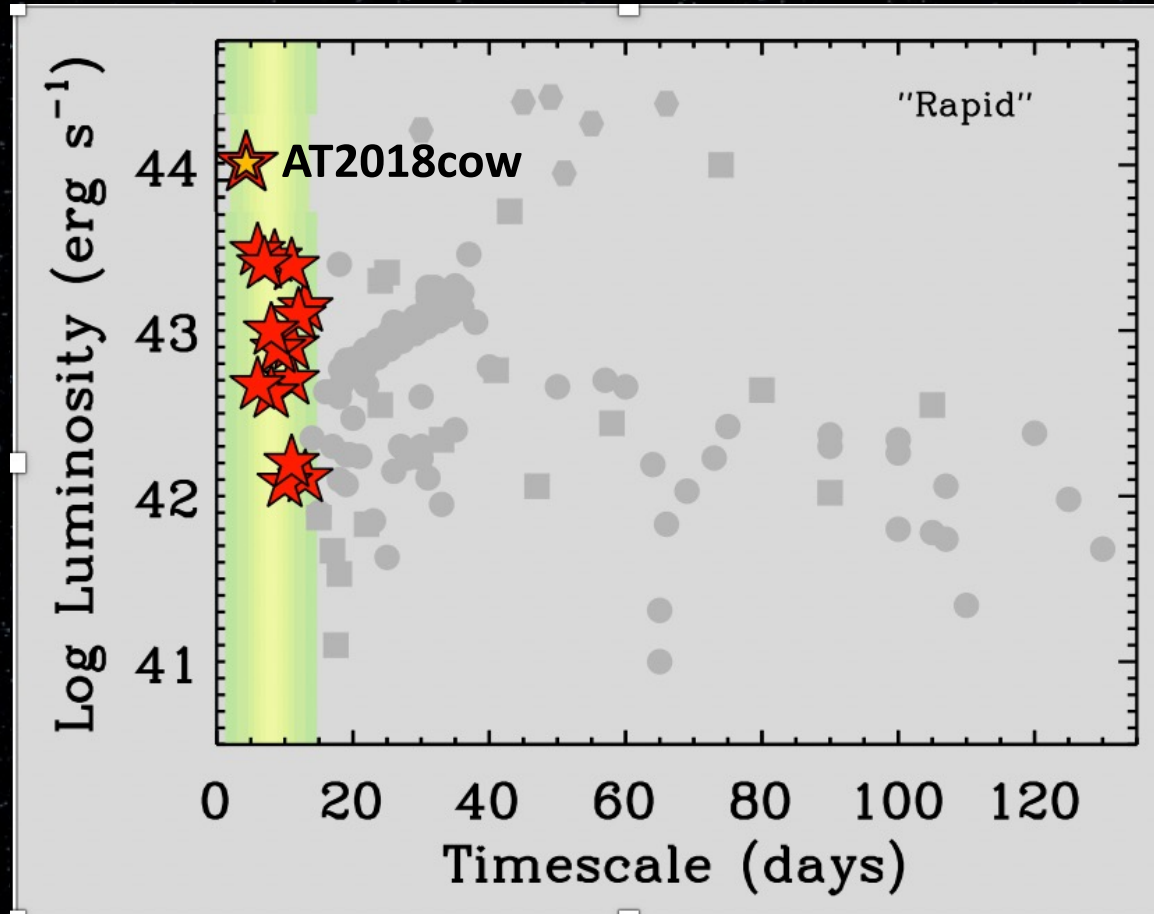
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Peculiar Transients: Science Themes

What is the origin of various classes of “peculiar” transients discovered in modern time domain surveys?

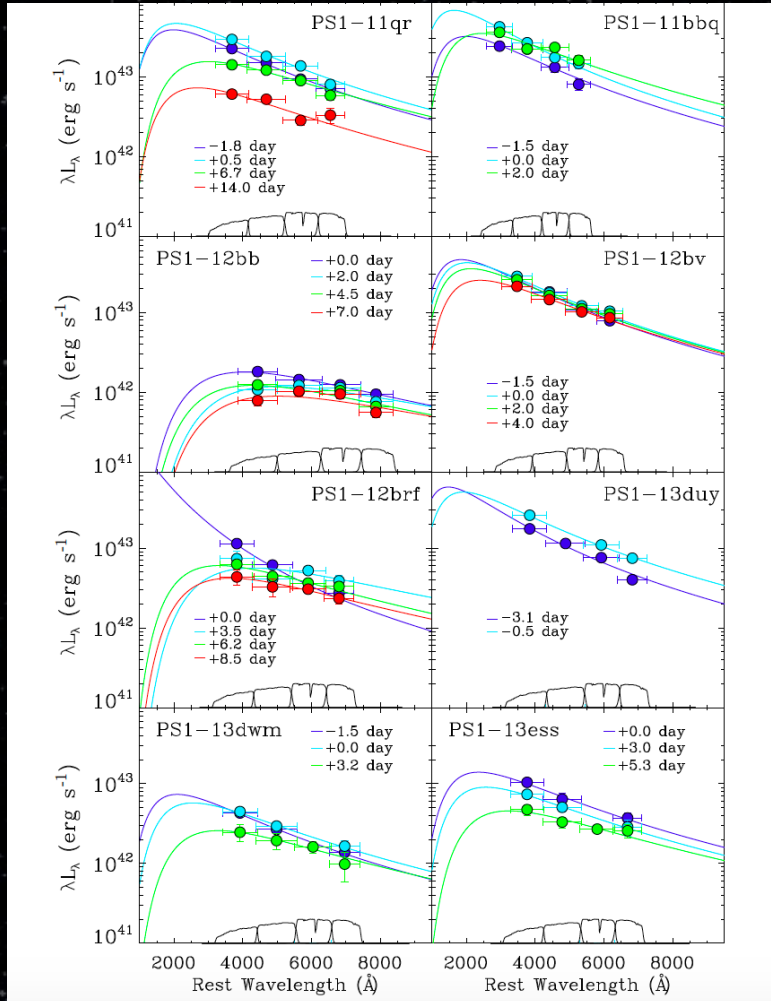
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Peculiar Transients: Science Themes



UVEX: A Time Domain Discovery Machine

Peculiar Transients: Science Themes



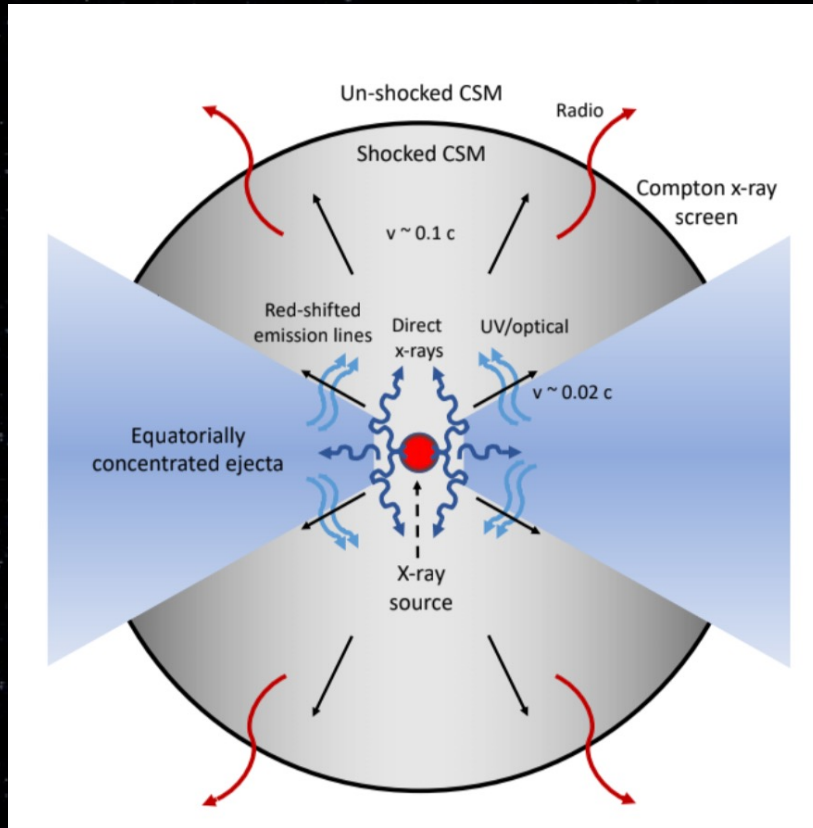
Many have:

hot temperatures
and colors that peak
in the UV at
maximum

Timescales and
luminosities
inconsistent with
being powered by
radioactive decay

UVEX: A Time Domain Discovery Machine

Peculiar Transients: Science Themes



Some some:
Evidence for central
engine activity
and/or circumstellar
interaction

Theories include:
Failed supernova,
mergers, tidal
disruption events, ...

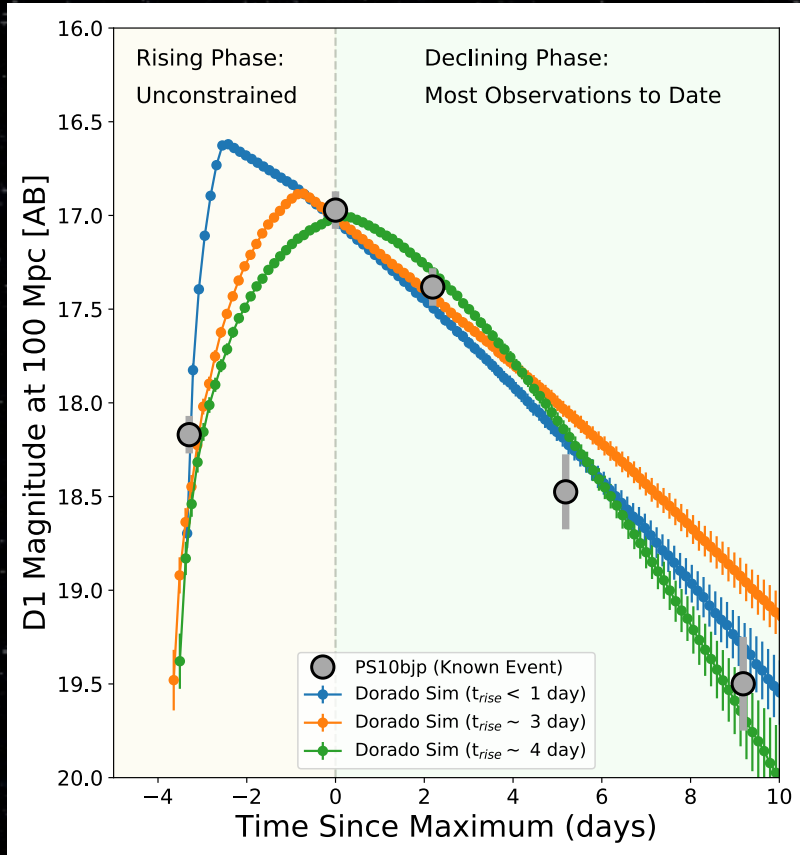
UVEX: A Time Domain Discovery Machine

Peculiar Transients: The Role and Potential of UV Data

1. Discovery

UVEX: A Time Domain Discovery Machine

Peculiar Transients: The Role and Potential of UV Data

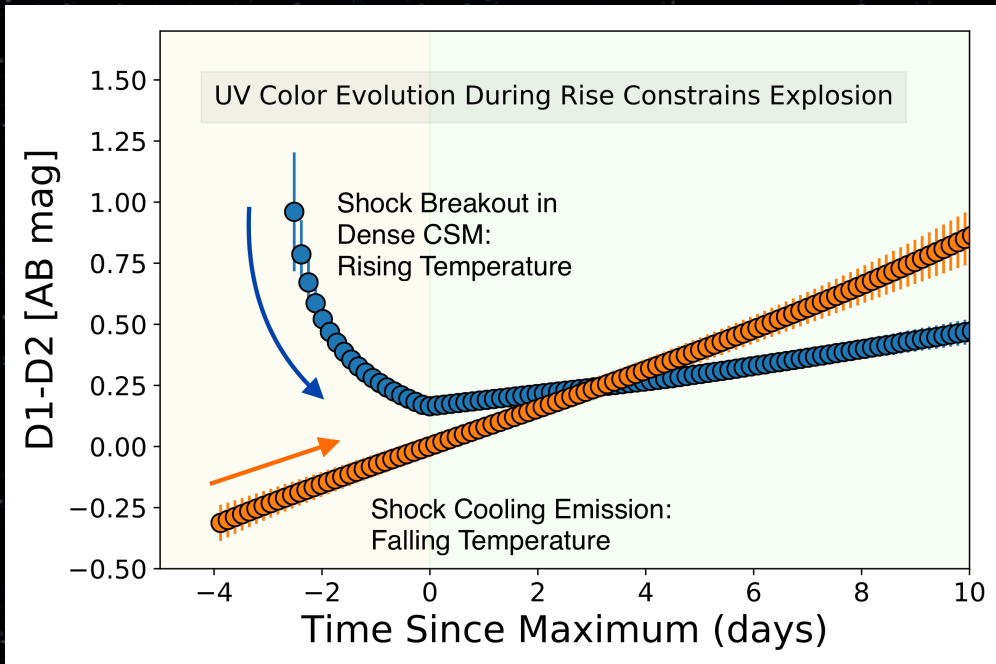


1. Discovery
2. Rise Time Measurements

UVEX: A Time Domain Discovery Machine

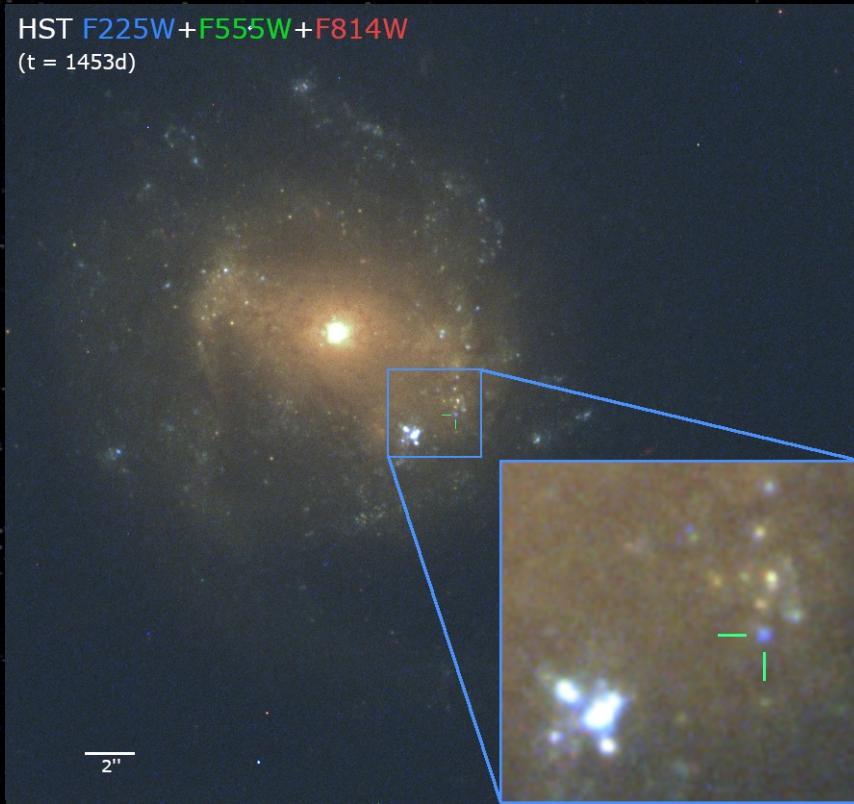
Peculiar Transients: The Role and Potential of UV Data

1. Discovery
2. Rise Time Measurements
3. Temperature Evolution on the rising phase



UVEX: A Time Domain Discovery Machine

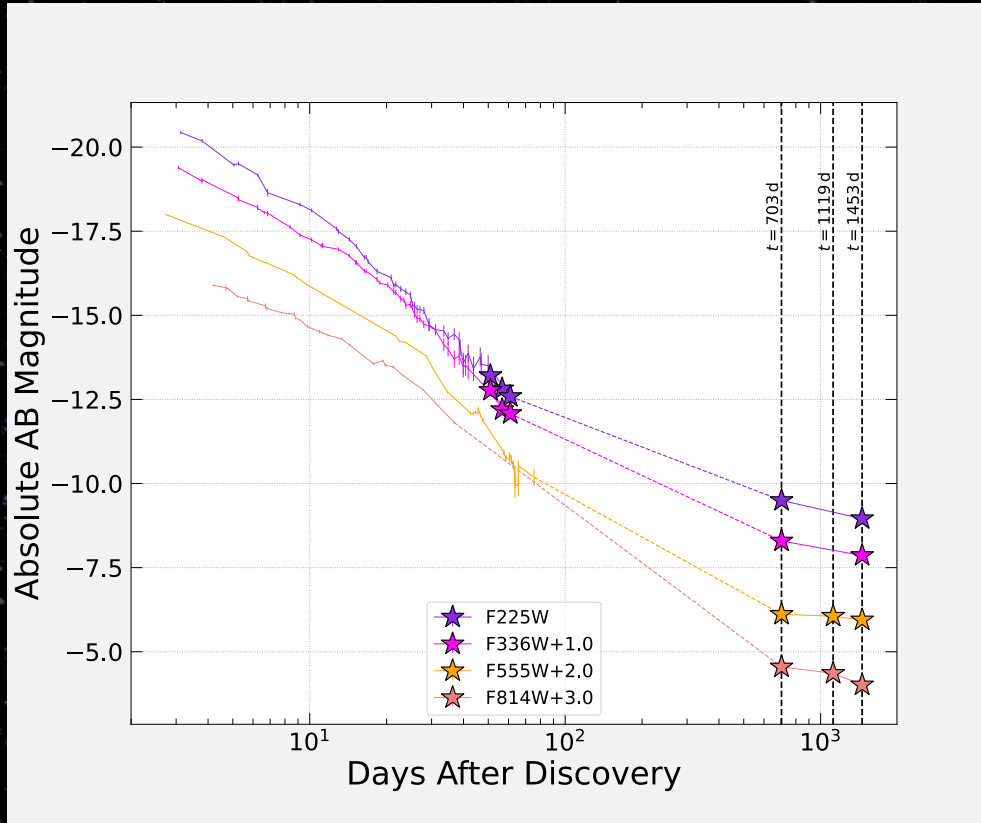
Peculiar Transients: The Role and Potential of UV Data



1. Discovery
2. Rise Time Measurements
3. Temperature Evolution on the rising phase
4. Probing the presence of a late-time time power source.

UVEX: A Time Domain Discovery Machine

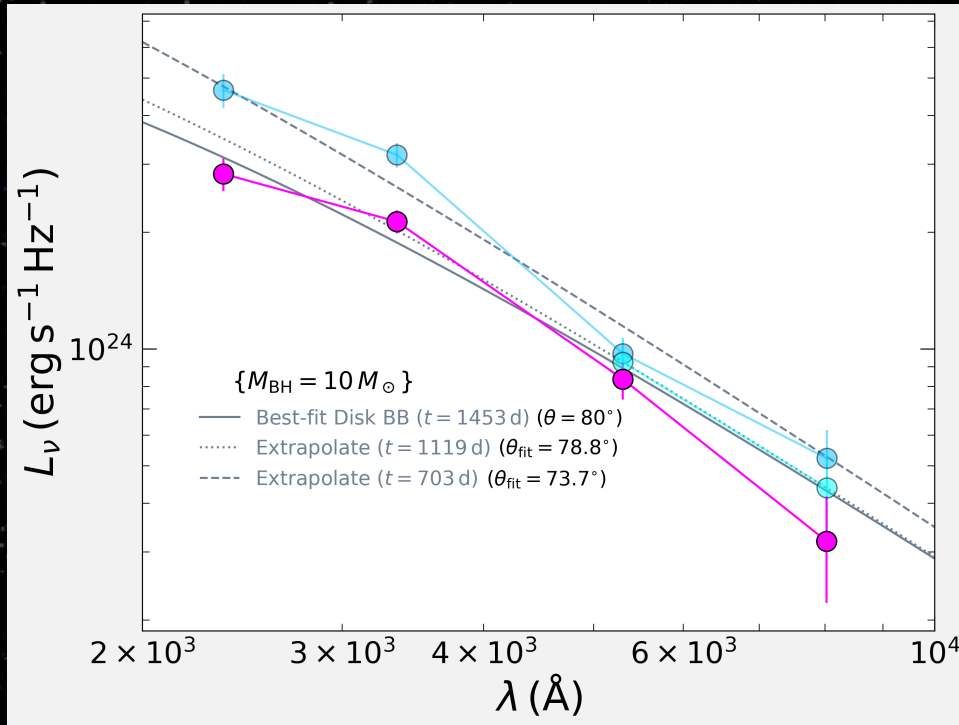
Peculiar Transients: The Role and Potential of UV Data



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Peculiar Transients: The Role and Potential of UV Data



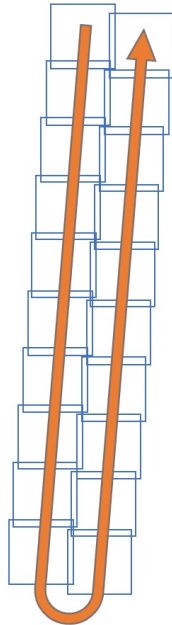
1. Discovery
2. Temperature Evolution on the rising phase
3. Probing the presence of a late-time power source.

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Multiwavelength Coordination will be critical

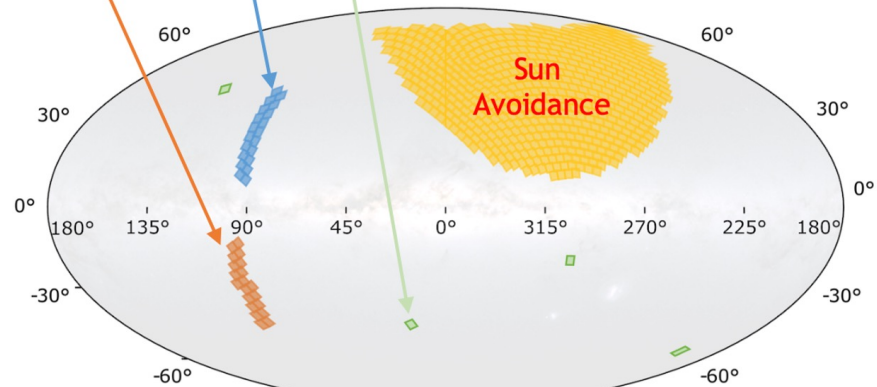
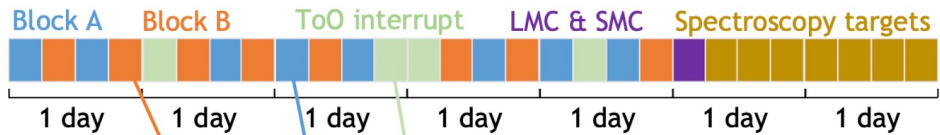
One 6-hr Survey Block =
20 contiguous fields
One dwell per field

One Dwell =
3 x NUV Images
1 x FUV Image
1 x LSS image
All simultaneous
w/overheads \approx 930s



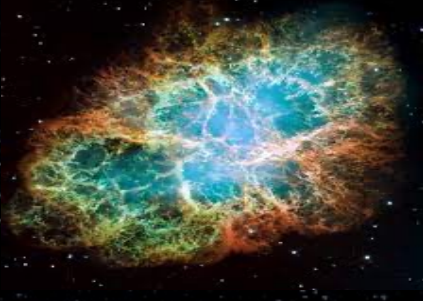
High Cadence Survey

2 blocks at \sim 12-hr cadence, including interrupts, \sim 5 days total

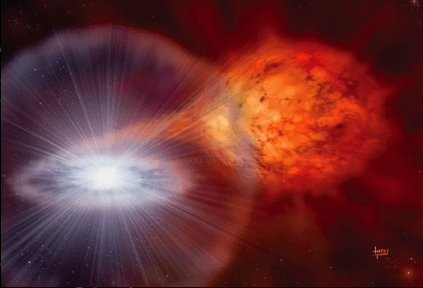


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